



Electro-optic modulator for Advanced LIGO

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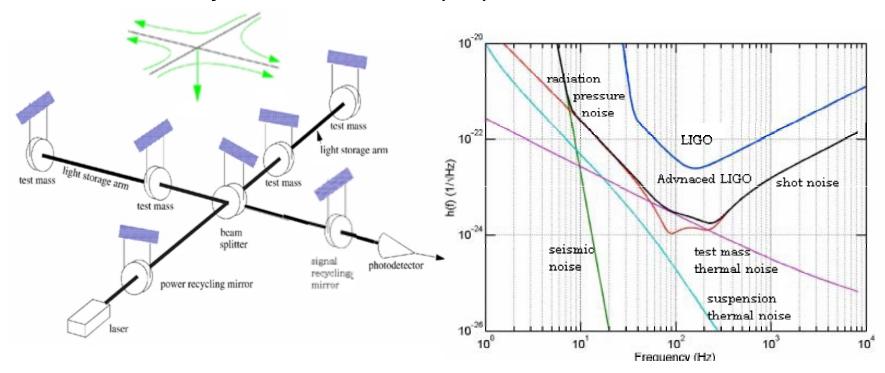
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Laser interferometer gravitational wave observatory (LIGO)

Michelson laser interferometer senses the small perturbation ($\sim 10^{-20} m/\sqrt{Hz}$) in the arms caused by Gravitational Waves (GW).

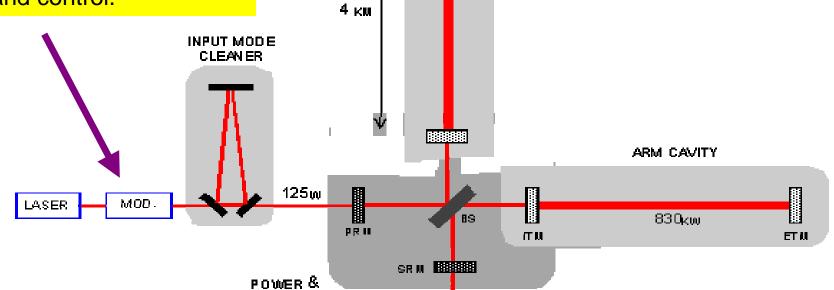






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Electro-optic modulator (EOM) modulate the input light, generate sidebands for length sensing and control.



OUTPUT MODE CLEANER

GW READOUT

EOM

SIGNAL RECYCLING

CAVITIES:

Problems:

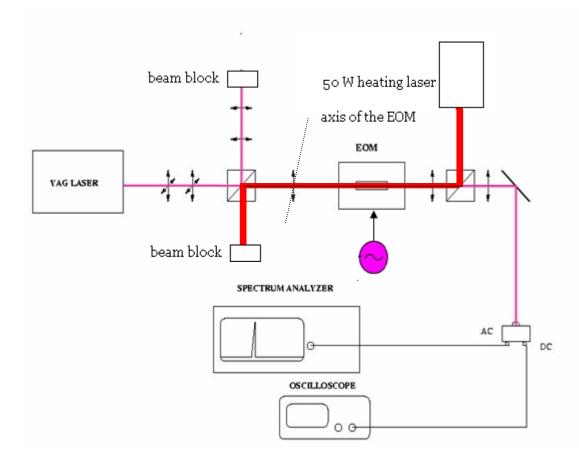
- 1. Residual intensity modulation,
- 2. Fluctuation of modulation index,
- 3. Phase noise of sidebands.



Residual intensity modulation measurement



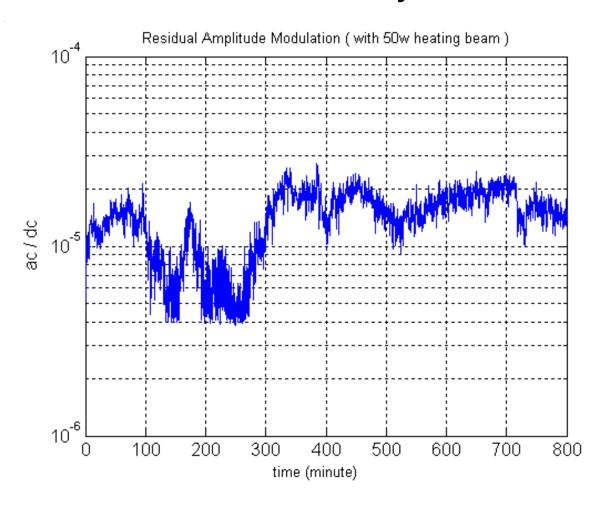
 Residual intensity modulation noise, which could be coupled into the readout signal of the interferometer, is due to misalignment of the EOM relative to the polarization direction of the incoming light.



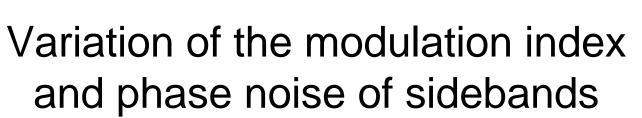




Residual intensity modulation



LIGO





$$E = E_0 e^{i[\varpi t + m \sin(\Omega t + \phi)]} \approx J_0(m) E_0 e^{i\varpi t} + J_1(m) E_0 e^{i[(\varpi + \Omega)t + \phi]} - J_1(m) E_0 e^{i[(\varpi - \Omega)t - \phi]}$$

Variation of the modulation index δm

Intensity fluctuation of the carrier light

Radiation pressure noise

For m = 0.1, δm must $< 10^{-7} / \sqrt{Hz}$

Phase noise $\delta \phi$ could be coupled into the power fluctuation of the signal in the dark port.

For RF sensing, the requirement on sidebands phase noise depends on the asymmetries of the interferometer caused by the recycling mirrors.

For example, at high frequencies, Issa must < -160 dBc/Hz.

For DC sensing, the coupling effect will be smaller then RF sensing. The requirement is relaxed, but to be determined.

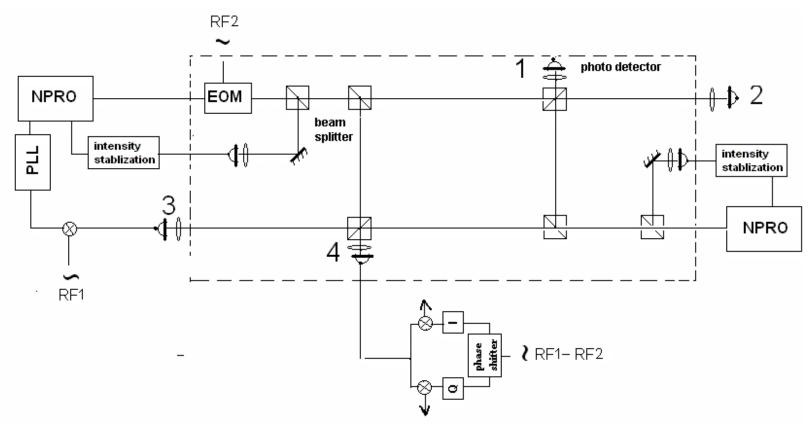




Experiment setup

• Phase lock two lasers, beat the sidebands of one laser with the reference laser and track the amplitude and phase information of the beat signal $(\delta_m < -\frac{160dB}{\sqrt{Hz}})$.

demodulated beat signal ~ $J_1(m)A_{\Omega}E_1E_2\cos(\Delta\phi)\sin(\delta\phi_{\Omega})$

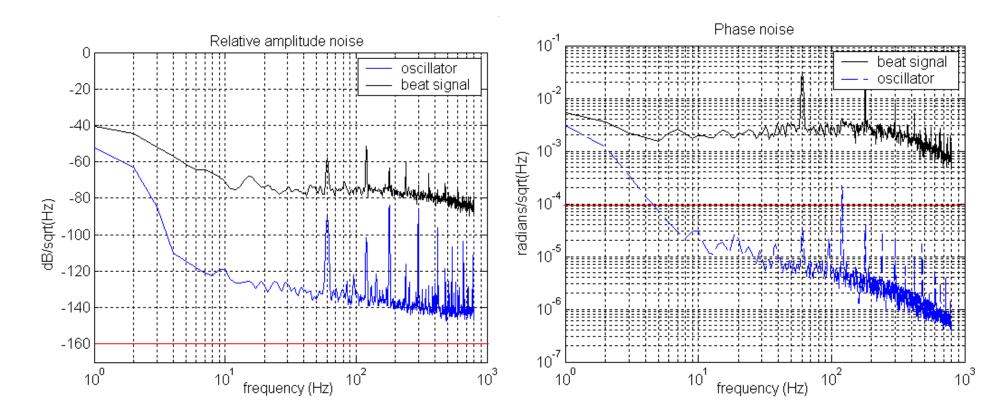






Progress

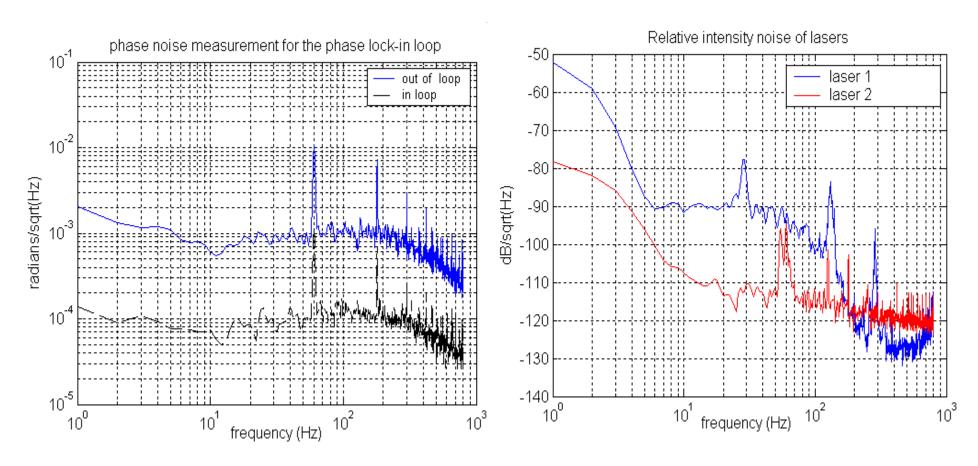
- Local oscillator characterization
- Signal from the photo detector without vacuum and intensity stabilization of the laser source







Phase locking noise and laser intensity noise







Need to be done

- Find or build better oscillators
- Intensity stabilization of laser sources
- Go to vacuum
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